# NASA TECHNICAL MEMORANDUM

NASA TM X-53484

June 29, 1966

ASA TM X-53484

GPO PRICE \$_	
CFSTI PRICE(S) \$_	
Hard copy (HC)	9.00
Microfiche (MF)	50
ff 653 July 65	

# THE ATMOSPHERE OF MARS: A DERIVATION OF ENGINEERING AND DESIGN PARAMETERS

By W. T. Roberts and George S. West Aero-Astrodynamics Laboratory

<b>#</b> 09	N66 31944	(THRU)
Y FORM	(PAGES)	(CODE)
FACILIT	TMX-53484	30 (CATEGORY)

NASA

George C. Marshall Space Flight Center, Huntsville, Alabama

### TECHNICAL MEMORANDUM X-53484

THE ATMOSPHERE OF MARS: A DERIVATION OF ENGINEERING AND DESIGN PARAMETERS

Ву

W. T. Roberts and George S. West

George C. Marshall Space Flight Center

Huntsville, Alabama

#### ABSTRACT

Three model atmospheres have been derived for use in mission planning, aerospace lander design, and Martian orbiter and flyby studies. Atmospheric parameters for these three models have been calculated from the planetary surface to 10,000 kilometers. These three model atmospheres of differing atmospheric composition, surface pressure, and surface temperature are the upper density model (60 percent  $CO_2$  and 40 percent  $N_2$ ); the mean density model (100 percent  $CO_2$ ); and the lower density model (80 percent  $CO_2$  and 20 percent Ar). The general program of Kern and Schilling [9], with a few revisions, was used; this publication contains the details of the mathematical basis and program routines in considerable depth. The models chosen should provide values which will be of use as guidelines for the engineering and design of orbiting and landing vehicles.

# NASA - GEORGE C. MARSHALL SPACE FLIGHT CENTER

Technical Memorandum X-53484

June 29, 1966

# THE ATMOSPHERE OF MARS: A DERIVATION OF ENGINEERING AND DESIGN PARAMETERS

Ву

W. T. Roberts and George S. West

SPACE ENVIRONMENT BRANCH AEROSPACE ENVIRONMENT DIVISION AERO-ASTRODYNAMICS LABORATORY

#### TECHNICAL MEMORANDUM X-53484

THE ATMOSPHERE OF MARS: A DERIVATION OF ENGINEERING AND DESIGN PARAMETERS

Вy

W. T. Roberts and George S. West

George C. Marshall Space Flight Center

Huntsville, Alabama

SUMMARY

31944

Three model atmospheres have been derived for use in mission planning, aerospace lander design, and Martian orbiter and flyby studies. Atmospheric parameters for these three models have been calculated from the planetary surface to 10,000 kilometers. These three model atmospheres of differing atmospheric composition, surface pressure, and surface temperature are the upper density model (60 percent CO<sub>2</sub> and 40 percent N<sub>2</sub>); the mean density model (100 percent CO<sub>2</sub>); and the lower density model (80 percent CO<sub>2</sub> and 20 percent Ar). The general program of Kern and Schilling [9], with a few revisions, was used; this publication contains the details of the mathematical basis and program routines in considerable depth. The models chosen should provide values which will be of use as guidelines for the engineering and design of orbiting and landing vehicles.

#### I. INTRODUCTION

The occultation experiment on Mariner IV has provided data that have permitted the formulation of more credible self-consistent model atmospheres for Mars. This experiment indicated pressures ranging from four to ten millibars surface pressure and from 150° to 270° Kelvin in surface temperature.

Mars is an obvious objective for planetary studies and exploration, and should be a major goal in the future space program. The first mission to Mars has been successfully flown in Mariner IV. This probe yielded a wealth of new information on surface conditions, the atmosphere, the planetary magnetic field, and the planetary mass. From these data, more realistic evaluations of the planetary conditions may be made. These new

evaluations will, in turn, be used in the design of more sophisticated instruments to narrow the range of values even more. In addition, new missions will be planned to gather extensive data on areas as yet untouched. These new missions will be in the form of unmanned orbiters and landers. These craft require extensive data on the magnetic field, gravitational field, and especially the atmospheric conditions of the planet. This report deals primarily with the atmosphere of the planet Mars, but we also touch on the gravitational and magnetic fields of Mars.

Ideally, one atmospheric model for the planetary conditions should be given; however, even for Earth this is not feasible since the upper atmosphere is subject to diurnal, latitudinal, and solar cycle variations. As a result, three atmospheric models were generated with the suggestion that (1) the mean density model be used for design, and (2) the design should be critically tested to verify its survival in either of the other two atmospheric models.

We wish to thank Mr. C. L. Hasseltine for the suggestions he has made during the construction of these models. We also wish to thank Mrs. Jeanette Scissum for revisions made in the original computer program.

### II. THE MARINER IV OCCULTATION EXPERIMENT

Probably the simplest experiment carried out by the recent Mariner IV probe has provided more data on the atmosphere of Mars than any other individual experiment. It consisted of observing the Doppler shift in signals from the probe as it was occulted by the planet Mars. This Doppler shift provided a direct indication of the atmospheric refractive index from which the surface pressure and temperature may be inferred. Johnson [2] interprets the low density near the surface of Mars as indicating that the atmosphere consists almost entirely of carbon dioxide.

Atmospheric composition from spectroscopic observations of Mars prior to Mariner IV had provided some data on the amount of carbon dioxide in the atmosphere. This was obtained from examining the pressure broadening of the weak  $\mathrm{CO}_2$  spectral absorption lines near 8700 A and broad  $\mathrm{CO}_2$  bands near  $2\mu$ . One of the prime candidates (before Mariner IV) for the major composition of the Martian atmosphere was thought to be molecular nitrogen. Nitrogen had not been observed in the spectral emissions of Mars; but since the spectral lines of nitrogen lie outside the earth's atmospheric windows, there was no problem in accounting for these missing data. Molecular nitrogen was selected because it was relatively heavy and would not readily escape from the planetary gravitational field. Also, molecular nitrogen does not have a high affinity for photochemical reaction. The large percentage of nitrogen in the terrestrial atmosphere undoubtedly played a role in the selection of this diatomic molecule as the primary gas of the Martian atmosphere.

Argon is now a primary candidate along with molecular nitrogen for a second constituent in the atmosphere of Mars. Argon results from the decay of potassium 40, and accounts for about 0.9 percent by volume of the earth's atmosphere. Owen [1] quotes a partial pressure of 2 millibars for argon on Mars using terrestrial analogy.

Polarization studies by Kaplan, Spinrad, and Munch [8] have indicated the presence of precipitable water vapor. But the amount of water vapor is almost indiscernible.

The molecular composition of the Martian atmosphere is now considered to be primarily carbon dioxide. For our mean atmospheric model of Mars, we chose a surface pressure of 8,000 dynes/cm² and a surface temperature of 210° K. An adiabatic lapse rate of 5°/km was chosen with the tropopause located at 14 kilometers. To this point, our model corresponds exactly to that given by Johnson [2]. However, at 60 kilometers, we assumed that our atmosphere would change from a purely mixed medium to a gas undergoing strong dissociation and diffusive equilibrium. When carbon dioxide is dissociated, the resulting atomic oxygen and carbon monoxide begin to undergo diffusive separation. The rate of change of mean molecular mass should be rather rapid since we assume that the predominant ion at the ionization peak is atomic oxygen, 0<sup>+</sup>. The mean molecular mass at 120 kilometers is then assumed to be 22, composed primarily of 60 percent oxygen atoms, 30 percent carbon monoxide, and 10 percent carbon dioxide.

A second point which we consider in developing the mean molecular mass curve has to do with the very low magnetic field which must surround Mars. It has been estimated that the surface magnetic field of Mars will not exceed about 500 gammas (5 x 10<sup>-3</sup> gauss). J. A. Van Allen et al. [6] indicate, as a result of Mariner IV data obtained from the University of Iowa "package" of low-energy-particle detectors aboard, that the equatorial surface magnetic field of Mars is less than 200 gammas, and that these results suggest interaction of the solar wind with the Martian atmosphere. This interaction may be of primary importance in determining the physical state of the atmosphere. J. J. O'Gallagher and J. A. Simpson [5] and E. J. Smith et al. [4] also comment on the solar wind interaction with the Martian atmosphere. If this is so, then the upper atmosphere of that planet must be related very intimately with the interplanetary environment. The solar wind will penetrate deeply into the upper atmosphere and probably will become an actual part of the atmosphere itself. How deeply the solar wind particles can pervade the atmosphere is a subject for further consideration. We feel that the upper atmosphere of Mars will definitely be influenced by the solar wind and that this will be even more significant when the upper atmospheric temperature is discussed. Essentially at 1,000 kilometers above the surface of Mars, the solar wind is the primary source of the atmospheric composition.

## ATMOSPHERIC TEMPERATURE

The temperature at the Martian surface probably varies from about  $150\,^\circ$  K to about  $270\,^\circ$  K. These two values make up our proposed temperature extremes. As we mentioned earlier, we have chosen an intermediate temperature of  $210\,^\circ$  K to represent our mean atmospheric surface temperature. On the other hand, once we exceed  $150\,^\circ$  kilometers, we have chosen to let our temperature vary by the equations

$$T = 85 \left(\frac{H}{150}\right)^{1.83635}$$
 for Model II and Model III

and

$$T = 92 \left(\frac{H}{150}\right)^{1.8175}$$
 for Model I

where H is the height in kilometers.

By this method, we reach the interplanetary temperature of about 190,000° K at 10,000 kilometers. This temperature distribution is based upon the supposition that the solar wind penetrates to low altitudes (relative to the earth) on Mars. The solar wind is cooled as it collides with the atmosphere, and in this manner becomes a part of the atmosphere. If one accepts this concept, then the atmospheric density and composition will be a direct function of solar activity. Also, there will be no isothermal exosphere on Mars, but rather a continuous increase in temperature until the kinetic gas temperature of interplanetary space is reached. We have chosen this altitude as 10,000 kilometers, although the solar wind predominates by 1,000 kilometers.

#### III. MODEL DESCRIPTIONS AND DATA OUTPUT TABLES

#### A. MODEL I (UPPER DENSITY MODEL)

Our upper density model supposes the atmosphere to be composed of 60 percent  $\mathrm{CO}_2$  and 40 percent  $\mathrm{N}_2$  by volume. We have chosen a surface temperature of 270° K which might well be a mean upper limit to the atmospheric surface temperature of the planet. Previous measurements of surface temperatures for the planet Mars undoubtedly were not indicators of atmospheric surface temperatures. Spectral measurements give the representative temperature of the planetary surface which may or may not be in equilibrium with the overlying atmosphere. This is especially true

on Mars where the tenuous atmosphere provides very poor thermal insulation. Planetary surface temperatures probably undergo a mean diurnal variation of as much as a factor of two greater than the mean diurnal variation of the atmospheric temperature immediately above the surface.

A dry adiabatic lapse rate of  $-4.55^{\circ}$  K/km exists to an altitude of twenty-two kilometers after which a lapse rate of  $-1.0^{\circ}$  K/km prevails to 100 kilometers. From 100 kilometers to 150 kilometers, the isothermal region exists; and from 150 kilometers to 10,000 kilometers, the temperature varies by the equation

$$T = 92 \left(\frac{H}{150}\right)^{1.8175}$$
.

The mean molecular mass is assumed to be 37.6, which exists to an altitude of 71 kilometers. At this altitude, dissociation is assumed to begin with the resulting components undergoing diffusive equilibrium. At the height of 120 kilometers, the predominant ion is taken as atomic oxygen and an  $F_2$ -type peaking of the ionosphere results.

The initial surface gravity is assumed as  $375 \text{ cm/sec}^2$  and varies with increasing altitude by the equation

$$g_z = 375 \left( \frac{r}{r + z} \right)^2$$

where  $\mathbf{g}_{\mathbf{z}}$  is the acceleration due to gravity at height,  $\mathbf{z}$ , above the surface.

The planetary radius r is taken as 3,381 kilometers. This equation for the acceleration due to gravity was used in all models discussed in this report.

The surface pressure, p, for this model was chosen as 10,000 dynes/cm² resulting in a calculated surface density of 1.67 x  $10^{-5}$  grams/cm³. This mass density was derived through the equation

$$\rho = \frac{Mp}{RT} ,$$

where  $\rho$  is the atmospheric mass density, M is the mean molecular mass of the atmosphere, T is the atmospheric gas temperature, and R is the universal gas constant. This equation was used throughout to calculate the value of atmospheric density.

The geopotential pressure scale height was calculated to be 15.92 kilometers at the surface dropping to 10.02 kilometers by 22 kilometers. The lowest calculated value of the potential pressure scale height was 7.13 kilometers at 71 kilometers.

The upper density model should also be used in total heating studies of vehicles entering into the Martian atmosphere. Apparently, a high argon content in the atmosphere will not increase heating significantly [11], which should be expected since argon is an inert gas and is not easily ionized.

We have found that the pressure and density curves for this model stabilize at relatively large values. The density at 1,000 kilometers is just over  $10^{-17}$  grams per cubic centimeter, whereas the pressure is reasonably stable at  $10^{-6}$  dynes per square centimeter which is the pressure at 700 kilometers on Earth.

Use of this model will result in atmospheric drag on aerospace vehicles; this drag is correspondingly less on Mars than on Earth below 700 kilometers and correspondingly greater on Mars than on Earth for altitudes greater than 700 kilometers.

At any rate, it is apparent in all models given here that the absence of an extensive isothermal region results in large scale heights above about 200 kilometers. This tends to stabilize the atmosphere of Mars more rapidly than on Earth.

Extensive data on the physical properties of the atmosphere may be found in Table I. In addition to the previously discussed parameters, we have calculated the speed of sound (cm/sec), columnar mass (grams/cm $^2$ ), number density (cm $^{-3}$ ), viscosity of the mixture (poise, or grams/(cm-sec) and kinematic viscosity (cm $^2$ /sec).

### B. MODEL II (MEAN DENSITY MODEL)

The mean density model was covered in Section II, but will be repeated here for continuity. An atmospheric temperature of 210° K was chosen as a mean at the surface of Mars. The pressure with this model is taken as  $8,000~\rm dynes/cm^2$ . A temperature lapse rate of  $-5^{\circ}/\rm km$  was chosen to obtain an altitude of fourteen kilometers after which a lapse rate of  $-0.64^{\circ}/\rm km$  obtains to 100 kilometers. Above 100 kilometers the atmosphere is isothermal to an altitude of 150 kilometers after which the temperature varies by

$$T = 85 \left(\frac{H}{150}\right)^{1.83635}$$
.

As in Model I, this temperature is a result of the interaction of the solar wind with the planetary atmosphere.

The mean molecular mass of the atmosphere was taken to be 44 at the surface. At an altitude of 60 kilometers, dissociation and diffusion began. By 60 kilometers, the atmospheric pressure is 1.06 dynes/cm<sup>2</sup> on Mars, whereas this pressure is not reached until about 95 kilometers on Earth. Since the bonding strength for CO2 is about the same as that for oxygen, we chose 60 kilometers as a reasonable mean altitude for photodissociation to begin. An interesting point here is that, if the mean molecular mass is not assumed to dissociate and the atmosphere is taken as isothermal above 100 kilometers, then interplanetary densities will be reached by 220 kilometers. This is not a reasonable model for the atmosphere of any terrestrial planet the size of Mars. Another subject for further consideration and possible interest might be the effect of the solar wind penetration of the atmosphere and the variations in the flux of the solar wind as a possible mechanism in accounting for the so-called "blue haze" phenomenon observed in the atmosphere of Mars.

The pressure scale height, using the initial parameters specified above, varies from 10.58 kilometers at the planetary surface to 6.5 kilometers at 30 kilometers. According to Mariner IV data, the pressure scale height should have been essentially a constant of about 9 kilometers in that region. For that to be true for any case, the temperature would have to be almost isothermal for the first 30 kilometers. An isothermal atmosphere here means that the atmosphere would be in radiative equilibrium. There is the possibility that the CO<sub>2</sub> abundance in the atmosphere maintains an isothermal balance by condensing out and revaporizing. A subadiabatic lapse rate could have been chosen assuming this or some similar mechanism acting in the lower atmosphere; however, we chose to go along with a near-dry adiabatic lapse rate because we need to know much more about the reactions taking place therein in order to construct better models. We may, indeed, speculate about these reactions but, without conclusive evidence, our speculations are only that.

In this model a density of about  $4.4 \times 10^{-22}$  is reached at 10,000 kilometers, which is on the order of the expected interplanetary density. The gravitational field of Mars at this point is still a strong influence on the atmospheric medium.

Table II is a tabulation of a number of parameters for the mean density model of the Martian atmosphere. All values are in cgs units.

### C. MODEL III (LOWER DENSITY MODEL)

This lower density model considers the atmosphere on Mars to be made up of 80 percent  $\mathrm{CO}_2$  and 20 percent argon. J. J. O'Gallagher and J. A. Simpson [5] report an implication based on their conclusions drawn from Mariner IV data that virtually all of the secondary production of particles from high-energy interactions takes place below the surface of Mars instead of in the atmosphere, as on Earth. Fast neutrons from cosmic radiation and solar flare protrons of sufficient energy will penetrate to the surface of Mars; the resulting  $\pi$  mesons produced in the solid planet will interact with nuclei below the Martian surface producing radioactive and stable isotopes.

From the above mechanism, we infer that the gases possibly produced and released, such as argon, may be in sufficient quantities to qualify as secondary gases in the Martian atmosphere.

The inclusion of argon in this model and nitrogen in Model I provides output data of interest in the consideration of aerodynamic heating factors in the Martian atmosphere in addition to the factors contributed by the 100 percent CO<sub>2</sub> atmosphere considered in Model II.

The surface pressure with this model is taken as 4,000 dynes/cm<sup>2</sup>. A temperature lapse rate of -5.15°/km was chosen to an altitude of ten kilometers after which a lapse rate of -0.15°/km obtains to 100 kilometers. Above 100 kilometers, the atmosphere is isothermal to an altitude of 150 kilometers, after which the temperature varies by

$$T = 85 \left(\frac{H}{150}\right)^{1.83635}$$
.

This minimum density model provides a probable lower limit for the atmosphere of Mars. Parachute descent to the Mars surface should certainly consider the effects of such an atmosphere as this.

Orbiters of the planet will certainly not encounter excessive drag for this model, since the atmospheric density at 100 kilometers on Mars is about the same as that at 300 kilometers on Earth. Above this altitude, the atmospheric density for Mars falls off at a much faster rate than it does on Earth. This is due to the relatively low temperatures between 100 and 200 kilometers. Once again, this atmosphere is found to stabilize by about one thousand kilometers because of the large temperatures encountered.

## Table 1

## Maximum Density Model

PLANET RADIUS= 3.38100000E 08 SUPFACE GRAVITY= 3.75000000E 07 MOLECULAR MASS= 3.76000000E 01

INITIAL CONDITIONS

MEIGHT: 0 PHESSUME: 1.00000000E 04 TEMPERATURE: 2.7000000E 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MAS' NUMBER DENSITY	TEMPERATURE HOLECULAR MASS DENSITY SCALE MEIGHT	DENSITY GRAVITY VISCOSITY(MIX.)
SPEED OF SOONS	MONBER DENSITY	DENSITY SOME HEIGHT	VISCUSITY(KIM.)
0	1.000000000 04	2.70000000E 02	1.67491854E-85
1.59211723E 06 2.89112703E 04	2.66666667E 01 2.68333771E 17	3.76000000E 01 1.59211723E 06	3.75000000E 02 1.98512356E 02
2107222	21000007722 17	11772117101 00	1.18520604E 07
5.00000000 <del>0</del> 05	7.20000000£ n3	2.47300000E 02	1.31663633F-05
1.45826145E 06 2.76692476E 04	1.92569142± 01 2.10934432± 17	3.76000000E 01	3.73691681E 02
2.700724706 84	5.78494495E 11	1.46258415E 06	2.07423218E 82 1.57540252E 8/
			1.979402522 07
1.00000000E 06	5.03000000E 03	2.24500000E 02	1.01323246E-05
1.32381600E 06	1.34930321E 01	3.76000000E 01	3.72785002E 02
2.63629171E 04	1.623269916 17	1.33168179E 06	2.17701416E 02 2.14858313E 07
			2.140303135 07
1.50000000E 06	3.36000000£ 03	2.01800000E 02	7.57448293E-06
1.16996021E 06	9.09384507E 00	3.7600000E 01	3.71679963E 02
2.49945817E 04	1.21348563t 17	1.20058955E 06	2.29619542E 02
			3.03148801F 07
2.00000000E 06	2.160000000 n3	1.79000000E 02	5.457053025-06
1.05551476E 06	5.82875499E 00	3.7600U000E 01	3.70576565E 02
2.35402899E 04	8.74258408£ 16	1.06811405E 06	2.43805170E 02
			4.46770756E 07
		. 4/=	
2.50000000E 06	1.320000000 03	1.66900000E 02	3.57663850E-06
9.84164320E 05 2.27307340E 04	3.57263872E 00 5.73002731E 16	3.76000000E 01 9.98881692E 05	3.69474807E 02 2.52488300E 02
2.2/30/3705 04	3.13005/3TC 70	7.700010725 03	7.059374316 0/

## Parameter Units

Height	centimeters	Temperature	degrees Kelvin
Pressure scale height .	centimeters	Molecular mass	dimensionless
Speed of sound	centimeters/sec	Density scale height	centimeters
Pressure	dynes/cm <sup>2</sup>	Density	grams/cm <sup>3</sup>
Columnar mass	particles/cm <sup>2</sup>	Gravity	cm/sec <sup>2</sup>
Number density	particles/cm <sup>3</sup>	<pre>Viscosity(mix) Viscosity(kinematic)</pre>	-

# Table ! (continued)

PLANET RADIUS = 3.38100000E 08 SURFACE GRAVITY = 3.75000000E 02 MOLECULAR MASS = 3.76000000E 01

INITIAL CONDITIONS

HEIGHT: 0 PRESSURE: 1.00000000E 04 TEMPERATURE: 2.70000000E 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND  3.00000000E 06 9.54680667E 05 2.23876607E 04	PRESSURE COLUMNAR MASS NUMBER DENSITY 7.850000000 n2 2.13098246E 00 3.51286860E 16	TEMPERATURE MOLECULAR MASS DENSITY SCALE HEIGHT  1.619000000 02 3.76000000E 01 9.71850839E 05	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.) 2.19270527E-06 3.68374689E 02 2.56357485E 02 1.16913790E 08
3.50000000E 06 9.25197015E 05 2.20392476E 04	4.61000000E n2 1.25518611E n0 2.12871284E 16	1.56900000E 02 3.76000000E 01 9.44653833E 05	1.32872600E-06 3.6/276211E 02 2.60410178E 02 1.95984558E 08
4.00000000E 06 8.95713362E 05 2.16852373E 04	2.66000000E n2 7.26419944E-01 1.26871175E 16	1.51900000E 02 3.76000000E 01 9.17289545E 05	7.91920008E-07 3.66179374E 02 2.64661360E 02 3.34202138E 08
4.50000000E 06 8.66229710E 05 2.13253511E 04	1.51000000E 02 4.13603245E-n1 7.44722112E 15	1.46900000E 02 3.76000000E 01 8.89756835E 05	4.64849753E-07 3.65084176E 02 2.69127780E 02 5.78956487E 08
5.00000000 06 8.36746057E 05 2.09592863E 04	8.39000000 01 2.30500446E-01 4.28369620E 15	1.41900000E 02 3.76000000E 01 8.62054555E 05	2.67384987E-07 3.63990620E 02 2.73828236E 02 1.02409727E 09
6.00000000 06 7.77778752E 05 2.02072721E 04	2.43000000E 01 6.71626148E-02 1.33475199E 15	1.31900000E 02 3.76000000E 01 8.06136648E 05	8.33141813E-08 3.61808427E 02 2.84018762E 02 3.40900861E 09
7.00000000 06 7.18811448E 05 1.94261682E 04	6.38000000E 00 1.77403176E-02 3.79189297E 14	1.21900000E 02 3.76000000E 01 7.49526453E 05	2.36687011E-08 3.59632/95E 02 2.95438830E 02 1.24822578E 10
8.00000000 06 7.29709993E 05 1.95728829F 04	1.60000000E no 4.47597867E-n3 1.03592662E 14	1.11900000€ 02 3.40000000€ 01 7.65507741€ 05	5.84707172E-09 3.57463724E 02 3.08357406F 02 5.2/370658E 10
9.00000000E 06 7.38337323E 05 1.96801810E 04	4.160000000E-01 1.17083754E-03 2.95772808E 13	1.01900000E 02 3.060000000E 01 7.79267311E 05	1.50248513F-09 3.55301213F-02 3.23133/60F-02 2.15066196F-11

PLANET RADIUS= 3.38190800E 08 SURFACE GRAVITY= 3.75080000E 02 HOLECULAR HASS= 3.76080800E 01

INITIAL CONDITIONS

MEIGHT= 0 PRESSURE: 1.0000000E 84 TEMPERATURE: 2.7000000E 02

MEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE MOLECULAR MASS DENSITY SCALE HEIGHT	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
1.0000000F 97 7.30315356E 05 1.95810000E 04	1.08000000E-n1 3.05823158t-n4 8.51426850t 12	9.19000000E 01 2.79000000E 01 7.75511626E 05	3.94350191E~10 3.53145265E 02 3.40260506E 02 8.62838699E 11
1.1000000E 07 8.18305158E 05 2.8/278405E 04	3.180000000=-^2 9.059935496-05 2.506979066 12	9.19000000E 01 2.4900000E 01 8.74267918E 05	1.n3628823E~10 3.50995877E 02 3.44260606E 02 3.28345529E 12
1.20000000E 07 9.3u401755E 05 2.21011520E 04	1.080000000±-n2 3.09585942E-n5 8.51426850± 11	9.19000000E 01 2.1900000E 01 1.00013647E 06	3.09543698E-11 3.44653J59E 02 3.40260606E 02 1.09923286F 13
1.30000000E 97 1.02399947E 06 2.31851778E 04	4.09000000E-03 1.17963715E-05 3.22438502E 11	9.19000000£ 01 1.99000000£ 01 1.10743428£ 06	1.00519834E-11 3.46716784E 02 3.40260606E 02 3.19434037E 13
1.4000000E 07 1.10738035E 06 2.41117126E 04	1.660000000£-n3 4.81735997£-n6 1.30867460£ 11	9.19000000E 01 1.8400000E 01 1.20511666E 06	3.997422095-12 3.445870795 02 3.402606065 02 8.512000946 13
1.50000000E 07 1.20566855E 06 2.51590140E 04	7.22000000±-r4 2.1082512U±-n6 5.69194617E 10	9.19000000E 01 1.6900000E 01 1.32021407E 06	1.59690102E-12 3.42463935E 02 3.40260606E 02 2.15075576E 14
1.60000000E 07 1.44185816E 06 2.75131883E 04	3.310000000t-n4 9.72535847E-07 2.31924459t 10	1.03400000E 02 1.5900000E 01 1.58866172E 06	6.12173023F-13 3.40347392E 02 3.20781384E 02 5.24004443E 14
1.70000000E 07 1.71067920E 06 3.00384191E 04	1.730000000±-04 5.114751826-07 1.08518359± 10	1.15500000E 02 1.4900000E 01 1.90548077E 06	2.684231676-13 3.382373306 02 3.035137946 02 1.138728766 15
1.80000000E 07 2.04171112E 06 3.27398585E 04	1.000000000±-04 2.97580517E-07 5.66016548± 79	1.280n0000E 02 1.3900000E 01 2.27778793E 06	1.30609401E-13 3.36133469E 02 2.88313144E 02 2.24744556E 15

PLANET RADIUS= 3.38100000E 08 SURFACE GRAVITY= 3.75000000E 82 MOLECULAR MASS= 3.76000000E 01

INITIAL CONDITIONS

HE!04T= 0 PRESSURE 1.00000000E 04 TEL/\_RATURE= 2.70000000E 02

HEIGHT	PRESSURE	TEMPERATURE	DENSITY
PRESSURE SCALE HEIGHT	COLUMNAR HASS	HOLECULAR HASS	GRAVITY
SPEED OF SOUND	NUMBER DENSITY	DENSITY SCALE HEIGHT	VISCOSITY(MIX.) VISCOSITY(KIM.)
1.9000000E 07	6.38000000E-05	1.41000000E 02	7.02038427E-14
2.42341910E 06	1.90996823E-07	1.29000000E 01	3.34036969E 02
3.56692448E 04	3.27823939E 09	2.72060355E 06	2.74700766E 02 3.91290214E 15
2.00000000E 07 2.88791137E 06	4.37000000E-05 1.31647669E-07	1.55000000E 02 1.19000000E 01	4.03521155E-14
3.89378154E 04	2.04262591E 09	3.26247254E 06	3.31946630E 02 2.62001380E 02
			6.49287841E 15
2.50000000E 07	1.200000008-05	2.33000000E 02	5.82267817E-15
5.49575282E 06	3.73142043E-08	9.40000000E 00	3.21593351E 02
5.37147115E 04	3.73133655E 08	6.40842636E 06	2.13693576E 02 3.67002210E 16
			0.070002102 10
3.00000000E 07	5.85000000E-06	3.24000000E 02	1.67213443E-15
9.32939345E 06 6.99852239E 04	1.87858800E-08 1.30812713E 08	7.70000000E 00 1.12346709E 07	3.11404098E 02 1.81216159E 02
	•		1.08374157E 17
3.50000000E 07	3.77000000E-06	4.29000000E 02	7.08155233E-16
1.41965107E 07	1.25091716E-08	6.70000000E 00	3.01378869E 02
8.63317329E 04	6.36682856E 87	1.76644485E 07	1.57485575E 02
			2.22388493E 17
4.00000000E 07	2.82000000E-06	5.47000000E 02	3.47231740E-16
2.16570063E 07 1.06629866E 05	9.67351323E-09 3.73508835E 07	5.60000000E 00 2.78589544E 07	2.91517666E 02 1.39468382E 02
			4.01657931E 17
4.50000000E 07	2.31000000E-06	6.77000000E 02	2.09297044E-16
2.94318538E 07	8.19670711E-09	5.10000000E 00	2.81820488E 02
1.24304961E 05	2.47207936E 07	3.91630333E 07	1.25364650E 02 5.98979552E 17
			7.40474778E 17
5.00000000E 07 3.95234771E 07	1.99000000E-06	8.20000000E 02	1.34266190E-16
1.44047997E 05	7.30845595E-09 1.75824067E 07	4.60000000E 00 5.44325863E 07	2.72287336E 02 1.13910170E 02
			8.48390579E 17
6.0000000E 07	1.62000000E-06	1.14000000€ 03	7.00750842E-17
6.16481600E 07	6.38516480E-09	4.10000000E 00	2.53713107E 02
1.79903541E 05	1.029554318 07	9.11189030E 07	9.66088698E 81 1.37864793E 18

PLANET RADIUS= 3.38100000€ 88 SUPFACE GRAVITY= 3.75000000€ 82 HOLECULAR MASS= 3.76000000€ 81

INITIAL CONDITIONS

HEIGHT= 0 PRESSURE= 1.00000000E 04 TEMPERATURE= 2.7000000E 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE MOLECULAR MASS DENSITY SCALE HEIGHT	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
7.0000000E 07 9.56550773E 07 2.24095773E 05	1.42000000E-n6 6.02218080E-n9 6.81318991E 06	1.51000000E 03 3.50000000E 00 1.52126453E 08	3.95866772E-17 2.35794980E 02 8.39423170E 01 2.12046088E 18
8.00000000E 07 1.29670890E 08 2.60916112E 05	1.300000000n6 5.94875958E-09 4.88005977E-06	1.93000000E 03 3.30000000E 00 2.22513736E 08	2.67343476E-17 2.18532953E 02 7.42490278E 01 2.7/728968E 18
9.0000000E 07 1.71651923E 08 3.00195369E 05	1.220000000E-n6 6.04178652E-09 3.68288100E 06	2.40000000E 03 3.1000000E 00 3.18775903E 08	1.89530841E-17 2.01927029E 02 6.65830686E 01 3.51304664E 18
1.0000000E 08 2.21717067E 08 3.41176582E 05	1.15000000E-n6 6.18355351E-09 2.87302192E 06	2.90000000E 03 2.9000000E 00 4.47065004E 08	1.38314416E-17 1.85977205E 02 6.05717427E 01 4.37928267E 18
2.00000000E n8 8.37597807E 08 6.63128079E 05	8.85000000E-07 1.41453920E-08 6.28611319E 05	1.02000000E 04 2.7000000E 00 5.02040226E 09	2.81758140E-18 6.25645440E 01 3.22975322E 01 1.14628569E 19
3.00000000E 08 1.79079169E 09 9.69621389E 05	8.11000000E-n7 1.70306011E-n7 2.79795456E 05	2.10000000E 04 2.6000000E 00 1.41021543E 11	1.207659546-18 4.76201630E 00 2.25091855E 01 1.86386448E 19
4.0000000E 08 3.32575600E 09 1.32137122E 06	7.77000000£-07 6.18157004E-08 1.563/1505E 05	3.60000000E 04 2.4000000E 00 9.92200466E 10	6.23016240E-19 1.25696222E 01 1.71916744E 01 2.75942636E 19
5.00000000E 08 5.44214610E 09 1.69030374E 06	7.580000000E-07 8.81524896E-09 1.01698499E 05	5.40000000E 04 2.2000000E 00 2.37337764E 10	3.71422094E-19 5.59873617E 01 1.40369434E 01 3.77924296E 19
6.00000000E D8 7.91846667E D9 2.03892006E 06	7.46n000005-07 3.31533108E-r9 7.20637175E n4	7.500000000 04 2.10000000E 00 1.31965509E 10	2.51227089E-19 2.25015235E 02 1.19107414E 01 4.74102590E 19

## Table ! (continued)

PLANET RADIUS= 3.38100000E 08 SURFACE GRAVITY= 3.75000000E 02 MOLECULAR MASS= 3.76000000E 01

INITIAL CONDITIONS

HEIGHT = 0 PRESSURE = 1.000000000 04 TEMPERATURE = 2.700000000 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE HOLECULAR MASS DENSITY SCALE HEIGHT	DENSITY GRAVITY VISCOSITY(HIX.) VISCOSITY(KIM.)
7.00000000 08 1.15526261E 10 2.46274820E 06	7.39000000E-n7 1.71999168E-09 5.40814518E 04	9.90000000E 04 1.90000000E 00 1.00830958E 10	1.70581706E-19 4.29653241E 02 1.03669697E 01 6.07742176E 19
8.00000000E 08 1.65635691E 10 2.94887670E 06	7.33000000E-07 1.04729040E-09 4.181569B1E 04	1.27000000E 05 1.7000000E 00 8.87459087E 09	1.18009993E-19 6.99901381E 02 9.15308650E 00 7.75619615E 19
9.00000000E 08 2.17559872E 10 3.37962916E 06	7.29000000E-07 7.03831238E-10 3.36408510E 04	1.57000000E 05 1.60000000E 00 7.87682272E 09	8.93547135E-20 1.03575946E 03 8.23226991E 00 9.21302255E 19
1.00000000E 09 3.00901733E 10 3.97458690E 06	7.26000000E-07 5.05139037E-10 2.76835714E 04	1.9000000E 05 1.4000000E 00 7.85109566E 09	6.43399418E-20 1.43722806E 03 7.48329086E 00 1.16308636E 20

END OF PROBLEM

## Table II

## Mean Density Model

INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 08 MOLECULAR MASS= 4.40000000E 01 SURFACE GRAVITY= 3.75000000E 02

INITIAL CONDITIONS

MEIGHT= 0 PRESSUME= 8.000000000 03 TEMPERATURE= 2.100000000 02

HEIGHT	PRESSURE	TEMPERATURE	DENSITY
PRESSURE SCALE HEIGHT	COLUMNAR MASS	HOLECULAR HASS	GRAVITY
SPEED OF SOUND	NUMBER DENSITY	DENSITY SCALE HEIGHT	VISCUSITY(MIX.)
SPEED OF SCOWN	MONBER DENSTIT	DEMOTITE SOURCE SECTIONS	VISCUSITY(KIM.)
_		2.10000000£ 02	2.01601137E-05
0	8.00000000E n3		3.75000000E 02
1.05819509E 06	2.1333333E n1	4.40000000E 01	
2.35701596E 04	2.76000450€ 17	1.05819509£ 06	2.09866256E 02
			1.040997386 07
5.00000000E 05	4.84000000E n3	1.85000000€ 02	1.38450943E-05
9.32219485E 05	1.29449256E D1	4.40000000E 81	3.73891681E 02
2.21227311E 04	1.89545174£ 17	9.34982843E 05	2:20104322E 02
2.2122/3115 04	1.003431.42 1.	,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,	1.58976398E 07
			8.99645078E-06
1.0000000E 06	2.72000000E 03	1.60000000E 02	
8.86243879E 05	7.29643088E 00	4.40000000E 01	.3.72785092E 02
2.05737220E 04	1.23165201£ 17	8.11034384E 05	2.32656654E 02
			2.58609380E 07
1.50000000E 06	1.39080000E 03	1.39400000E 02	5.27684468E-06
7.82439979E 05	3.73977652£ 00	4.40000000E 01	3.71679963E 02
1.920367125 04	7.22422268£ 16	7.08714535E 05	2.45446391E 02
1.72030/125 04	7.224222000 10	7.007147032 05	4.65138554E 07
			4.052005542 01
		. 7.000000 00	2.62658750E-06
2.0000000E 06	6.76000000t n2	1.36200000E 02	
6.86315102E 05	1.82418443E 00	4.40000000£ 01	3.70576565E 02
1.89819764E 04	3.59590895£ 16	6.94507391E D5	2.47689508E 02
			9.43008781E 07
•			
2.50000000E 06	3.23000000E n2	1.33000000E 02	1.285207256-06
6.70190224E D5	8.74213868E-01	4.400000000 01	3.69474807E 02
1.87576616E 04	1.75950287E 16	6.80212371E 05	2.50013100E 02
1.8/2/00105 04	1./595020/2 10	2.405150.15 43	1.94531349E 88
			1.7-2313-75 00

## Parameter Units

Heightcentimeters	Temperature	degrees Kelvin
Pressure scale height centimeters	Molecular mass	dimensionless
Speed of sound centimeters/sec	Density scale height .	centimeters
Pressure dynes/cm <sup>2</sup>	Density	
Columnar mass particles/cm <sup>2</sup>	Gravity	cm/sec <sup>2</sup>
Number density particles/cm <sup>3</sup>	<pre>Viscosity(mix) Viscosity(kinematic) .</pre>	poise cm <sup>2</sup> /sec 15

INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 08 SURFACE GRAVITY= 3.79000000E 02 MOLECULAR MASS= 4.40000000E 01

INITIAL CONDITIONS

MEJOHT= 0 PRESSURE= 8.000000000 03 TEMPERATURE= 2.10000000 02

HEIGHT. Pressure scale Height Speed of Sound	PRESSURE TEMPERATURE COLUMNAR MASS MOLECULAR MASS NUMBER DENSITY DENSITY SCALE MEIGHT	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
3.0000000E 06 6.54065347E 05 1.85306316E 04	1.520000000 n2 1.29800000E 02 4.12623355E-01 4.40000000E 01 8.48414326E 15 6.65828876E 05	6.19713818F-07 3.68374689E 02 2.52422098E 02 4.07320428E 08
3.50000000F 06 6.3/940469E 05 1.83007854E 04	7.00000000000 01 1.266000000 02 1.905922526-01 4.400000000 01 4.005930706 15 6.513563056 05	2.92608285E-07 3.6/276211E 02 2.54921864E 02 8.71205215E 08
4.00000000E 06 6.21815592E 05 1.80680155E 04	3.160000000E 01 1.234000000E 02 8.62965046E-02 4.40000000E 01 1.85528665t 15 6.36794051E 05	1.35517134E-07 3.66179374E 02 2.57518246E 02 1.90026338E 09
4.50000000E 06 6.05489153F 05 1.78292402F 04	1.400000000£ 01 1.20160000€ 02 3.834/3207E-02 4.40000000€ 01 8.44125877E 14 6.21934466€ 05	6.16581373E-08 3.65084176E 02 2.60252056E 02 4.22088742E 09
5.00000000E 06 5.89565836E 05 1.75932392E 04	6.060000006 n0 1.170000006 02 1.66487807E-n2 4.40000000E 01 3.752544586 14 6.073980396 05	2.74100007E-08 3.63990620E 02 2.63027020E 02 9.59602382E 09
6.00000000E 06 5.57316081E 05 1.71052899E 04	1.060000006 00 1.10600000E 02 2.92972723E-03 4.40000000E 01 6.943682206 13 5.77635884E 05	5.07192734E-09 3.61808427E 02 2.69007286E 02 5.30384/39E 10
7.0000000E 06 6.06379810E 05 1.78422897E 04	1.94000000E-01 1.04200000E 02 5.39439125E-04 3.81000000E 01 1.34887936E 13 6.32286410E 05	8.93196289E-10 3.59632795E 02 2.79930382E 02 3.22994173F 11
4.0000000F 06 6.35892350E 05 1.82713843E 04	3.90000000£-02 9.78000000£ 01 1.09101980£-04 3.41000000£ 01 2.88911514£ 12 6.67087638£ 05	1.63549695E-10 3.57463724F 02 2.82683736E 02 1.74842717E 12
9.0000000E 06 6.60093742F 05 1.8615860WE 04	8.44000000E-03 9.14000000E 01 2.37544925E-05 3.0700000E 01 6.69014220E 11 6.966931196 05	3,40960630E-11 3,59301213E 02 2,94979448E 02 8,92226401E 12

### INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 88 SURFACE GRAVITY= 3.75000000E 82 MDLECULAR MASS= 4.40000000E 81

INITIAL CONDITIONS

HEIGHT= 0 PRESSURE= 8.0000000E 03 TEMPERATURE= 2.10000000E 02

MEIBHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE MOLECULAR MASS DENSITY SCALE HEIGHT	DENSITY GRAVITY VISCOSITY(MIX.)
1.00000000	1.910000000E-03 5.49853918E-06 1.62799677E 11	8.50000000E D1 2.80000000E D1 7.14723233E D5	VISCUSITY(KIM.) 7.56731968E-12 3.53145265E 02 2.99343308E 02 3.95573758E 13
1.1000000E 07 7.53838027E 05 1.98938424E 04	5.060000000000004-04 1.44161238F-06 4.31291291E 10	8.50000000£ 01 2.50000000E 01 8.05391969£ 05	1.78995127E-12 3.50995877E 02 2.99343308E 02 1.67235451E 14
1.20000000E 07 8.50634121E 05 2.12069072E 84	1.57000000E-04 4.50046231E-07 1.33819630E 10	8.50000000E D1 2.20000000E D1 9.20839866E D5	4.88734521E-13 3.48853050E 02 2.9934350E 02 6.12486525E 14
1.3000000E 07 9.42297533E 05 2.22419919E 04	5.44000000E-05 1.56900394E-07 4.63680756E 09	8.50000000E 01 2.00000000E 01 1.01916490E 06	1.53949959E-13 3.46716784E 02 2.99343308E 02 1.94441954E 15
1.40000000E 07 1.01870004E 06 2.31261220E 04	2.04000000E-n5 5.92012912E-n8 1.73880283E 09	8.50000000E 01 1.85000000E 01 1.10860951E 06	5.34013920E-14 3.44587079E 02 2.99343308E 02 5.60553381E 15
1.5000000E 07 1.10856533E 06 2.41248275E 04	8.27000000E-06 2.41465282t-08 7.04897031E 08	8.50000000E 01 1.7009000E 01 1.21390739E 06	1.98932213E-14 3.42463935E 02 2.99343308E 02 1.50475031E 16
1.60000000E 07 1.17787192E 06 2.48673030E 04	3.54000000E-R6 1.04011387E-08 3.01733433£ 08	8.50000000E 01 1.6000000E 01 1.29779758E 06	8.01445375E-15 3.40347352E 02 2.99343308E 02 3.73504318E 16
1.7000000E 07 1.25639671E 06 2.56828401E 04	1.74000000E-n6 5.14431686E-n9 1.48309654E n8	8.50000000E 01 1.5000000E 01 1.39295319E 06	3.69310104E-15 3.38237330E 02 2.9934330BE 02 8.10547302E 16
1.80000000E 07 1.88142768E 06 3.14264828E 04	9.66000000t-07 2.8/385500£-n9 5.89114597£ n7	1.18800000E 02 1.4000000E 01 2.09897141E 06	1.36917301E-15 3.36133869E 02 2.61432768E 02 1.90942099E 1/
1.9000000E 07 2.23763686E 06 3.42747626E 04	5.9000000E-07 1.76627156E-09 3.25804647E-07	1.31200000E 02 1.3000000E 01 2.51203878E 06	7.03122729E-16 3.34036969E 02 2.51357319F 02 3.57487120E 1/

### INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 08 SURFACE GRAVITY= 3.75000000E 82 MOLECULAR MASS= 4.40000000E 01

INITIAL CONDITIONS

ME|8HT= 0 PRESSURE= 8.000000000 83 TEMPERATURE= 2.10000000 02

MEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE MOLECULAR MASS DENSITY SCALE MEIGHT	AISCOSITA(WIN") AISCOSITA(WIN") AUSCOSITA(WIN")
2,00000000E 07 2,66430008E 06 3,73999672E 04	3.900000000E-07 1.17488766E-09 1.95946921E 07	1.44200000E 02 1.20000000E 01 3.00985894E 06	3.90346420E-16 3.31946630E 02 2.42222770E 02 6.20532834E 1/
2.50000000E 07 5.06915230E 06 5.158/8373E 04	9.58000000E-08 2.97891731E-10 3.19554388E 06	2.17200000E 02 9.50000000E 00 5.91098077E 06	5.03963289E-17 3.21593351E 02 2.07250803E 02 4.11241865E 18
3.00000000E 07 8.62706791E 06 6.72994105E 04	4.42000000E-08 1.41937760E-10 1.05512198E 06	3.03500000E 02 7.80000000E 00 1.03889143E 07	1.36624248E-17 3.11404098E 02 1.83549688E 02 1.34346349E 19
3.50000000E 07 1.31367362E 07 8.30468934E 04	2.74000000E-08 9.09154649E-11 4.92711153E 05	4.02900000E 02 6.80000000E 00 1.63457912E 07	5.56201065E-18 3.01378869E 02 1.66357933E 02 2.99096754E 19
4.00000000E 07 1.99934337E 07 1.02452685E 05	2.010000006-n8 6.89495092E-11 2.83316610E 05	5.14000000E 02 5.70000000E 00 2.57189821E 07	2.68088017E-18 2.91517666E 02 1.53406181E 02 5.72223192E 19
4.50000000E 07 2.72456165E 07 1.19599116E 05	1.62000000E-n8 5.74834005E-11 1.83676356E 05	6.39000000E 02 5.20000000E 00 3.62539510E 07	1.58557616E-18 2.818204886 02 1.43091437E 02 9.02457039E 19
5.00000000E 07 3.66069029E 07 1.38631252E 05	1.38000000E-n8 5.06817548E-11 1.28841705E n5	7.76000000E 02 4.70000000E 00 5.04158174E 07	1.00527480E-18 2.72287336E 02 1.34786011E 82 1.34081348E 20
6.00000000E 07 5.70129600E 07 1.73008104E 05	1.10000000E-08 4.33560573E-11 7.37917869E 14	1.08000000E 03 4.2000000E 00 8.42678577E 07	5.14502901E-19 2.53713107E 02 1.22307307E 02 2.37874803E 20
7.00000000E 07 8.00000207E 07 2.15779017E 05	9.58000000E-n9 4.06285156E-11 4.81994536E-04	1.44000000E 03 3.60000000E 00 1.41044394E 08	2.88034604E-19 2.39794980F 82 1.13142409F 82 3.92782492E 20

## Table 11 (continued)

### IMPUT CONSTANTS

PLAMET RADIUS= 3.38188888E 88 SUMFACE GRAVITY= 3.75888898E 82 HOLECULAR MASS= 4.4888888E 81

INITIAL COMBITIONS

MEIGHT= 0 PRESSURE= 8.000000000 03 TEMPERATURE= 2.100000000 02

MEIGHT PRESSÜRE SCALE MEIGHT SPEED OF SOUND	NUMBER DENSITY	TEMPERATURE MOLECULAR MASS DENSITY SCALE MEIGHT	VISCOSITY(KIM.)
8.0000000E 07 1.1998000E 08 2.50985520E 05	8.69000000E-n9 3.97651698E-11 3.42169308E 04	1.8400000E 83 3.4000000E 80 2.85898111E 88	1.03130320E-19 2.10532053E 02 1.00250032E 02 5.50146799E 20
9.00000000E 07 1.57973410E 08 2.87986181E 05	8.07000000E-n9 3.99649321E-11 2.56435286E n4	2.28000000E 03 3.2000000E 00 2.93373449E 08	1.36225457E-19 2.01927029E 82 1.00875956E 82 7.48507388E 20
1.00000000 08 2.04718756E 08 3.27837381E 05	7.63000000E-n9 4.10265333E-11 1.99564766E 04	2.77000000E 03 3.00000000E 00 4.12790020E 08	9.93883845E-20 1.85977205E 02 9.64713187E 01 9.70649832E 20
2.0000000E 08 7.83136353E 08 6.41207132E 05	5.73000000£-n9 9.15854194E-11 4.19756498E 03	9.89000000E 03 2.80000000E 00 4.69397064E 09	1.95112894E-20 6.25645440E 01 7.61679006E 01 3.90481117E 21
3.00000000E 08 1.70804259E 09 9.46954254E 05	5.22000000E-n9 1.09617432E-09 1.81821931± 03	2.08000000E 04 2.7000000E 00 1.34505203E 11	8.14967970E-21 4.76201630E 00 6.91090765E 01 8.47997456E 21
4.00000000E 08 3.13064499E 09 1.28202520E 06	4.99000000E-09 3.96988861E-10 1.02415323E 03	3.53000000E 04 2.50000000E 00 9.33991372E 10	4.25045534E-21 1.25696222E 01 6.54552933E 01 1.53995956E 22
5.00000000E 08 5.12841215E 09 1.64085843E 06	4.86080000E-09 5.65199339E-11 6.61856342E 02	5.32000000E 04 2.30000000E 00 2.23655491E 10	2.52709798E-21 8.59073617E 01 6.32173558E 01 2.50157910E 22
6.000000000 08 7.49806807E 09 1.98405790E 06	6.78000000E-09 3.90195802E-11 8.54989297E 02	7.44000000E 04 2.2000000E 00 1.24959340E 10	3 122582126-21 2.25015235E 02 6.16994775E 81 1.97591208E 22
7.00000000 08 1.09417372E 10 2.396/5031E 06	4.73000000±-n9 1.10088777±-11 3.47202694± n2	9.87000000E 04 2.00000000E 00 9.54991391E 09	1.15277246E-21 4.29653241E 82 6.06038968E 81 5.25722469E 22
8.00000000E 08 1.55201947E 10 2.8544800E 06	4.69000000E-09 6.70094405E-12 2.69675440E 02	1.26000008E 05 1.80000000E 07 8.31556153E 09	6.05031017E-22 6.99901301E 02 5.97742797E 01 7.41771154E 22

#### INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 08 SURFACE GRAVITY= 3.75000000E 02 MOLECULAR MASS= 4.40000000E 01

INITIAL CONDITIONS

MEIGHT: 0 PRESSURE: 8.00000000E 03\_ TEMPERATURE: 2.1000000E 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE Molecular mass Density scale height	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
9.00000000E 08 2.04762232E 10 3.27872188E 06	4.67000000E-09 4.50876801E-12 2.15504491E 02	1.57000000E 05 1.70000000E 00 7.41348021E 09	6.08185074E-22 1.03575966E 03 5.91089423E 01 9.71890709E 22
1.00000000E 09 2.80841618E 10 3.83981574E 06	4.65000000E-n9 3.23539466E-12 1.77312131E 02	1.90000000E 05 1.50000000E 00 7.32768928E 09	4.41530002E-22 1.43722806E 03 5.85882641E 01 1.32693733E 23

## Table III

## Minimum Density Model

INPUT CONSTANTS

PLANET RADIUS= 3.38199898 98 SURFACE GRAVITY= 3.78098998 92 NDLECULAR MASS= 4.32998998 91

INITIAL CONDITIONS

ME:8MT= 8 PHESSURE= 4.80088808E 03 TEMPERATURE= 1.50080008E 02

MEISHT PRESSURE SCALE HEISHT SPEED OF SOUND 7.69859926E 85 2.91349229E 84	PRESSURE COLUMNAR MASS NUMBER DENSITY 4.08000300E 03 1.0666667E 01 1.93200319E 17	TEMPERATURE MOLECULAR MASS DEMESTY SCALE MEIGHT 1.900000000 02 4.320000000 01 7.69650926E 05	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(MIM.) 1.305549636-05 3.75000000000000000000000000000000000000
9.0000000E 89 6.36410099E 05 1.82788211E 84	1.960000000 03 9.24219996E 00 1.14517929E 17	1.24000000 02 4.32000000 01 6.30296596E 05	0.212733706-06 3.73001001E 02 6.32379121E 02 7.69993496E 07
1.0000000E 06 5.05935441E 05 1.62912893E 04	8.15000000E 02 2.18624675E 00 5.99460368E 16	9.89000008 01 4.380000008 01 9.089398108 05	4.29907214E-06 3.72705002E 02 7.09524073E 02 1.65041351E 08
1.9000000E 06 9.01942804E 09 1.62332982E 04	3.020000006 02 8.125269856-01 2.237212246 16	9.78000000E 01 4.3200000E 01 9.06426415E 05	1.604432478-06 3.716799438 02 7.120593456 02 4.438076136 08
2.00000000 06 4.978365326 05 1.61667681E 04	1.11000000€ 02 2.99533242E-01 8.29066362E 19	9.70000000£ 01 4.32000000£ 01 9.03779427£ 05	5.94572208E-07 3.70576569E 02 7.14989643E 02 1.20252787E 09
2.50000000 06 4.94244294E 05 1.61063267E 04	4.05000000E 01 1.09615092E-01 3.04696758E 15	9.6300000E 01 4/3200000E 01 5.01635313E 05	2.185154216-07 3.694748076 02 7.175635506 02 3.263903476 09

## Parameter Units

Height centimeters Pressure scale height centimeters Speed of sound centimeters/sec Pressure dynes/cm <sup>2</sup> Columnar mass particles/cm <sup>2</sup> Number density particles/cm <sup>3</sup>	Temperature	dimensionless centimeters grams/cm <sup>3</sup> cm/sec <sup>2</sup>
Number density particles/cm <sup>2</sup>	Viscosity (mix) Viscosity (kinematic)	

INPUT CONSTANTS

FLANET RADIUS: 3.36100000E 08 SUFFACE GRAVITY: 3.75000000E 02 MOLECULAM MASS: 4.32000000E 01

INITIAL CONDITIONS

MEIGHT= 0 PHESSURE: 4.000000006 03 TEMPERATURE: 1.500000006 08

HETGHT PHESSURE SCALE HE SPEED OF SOUN	PRESSURE 1987 COLUMNAR MASS D. Number DENSITY	TEMPERATURE HOLFCULAR HASS DENSITY SCALE HEIGHT	DENSITY  URAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
3.00000000E 0 4.90138423E 0 1.60412802E 0	3.99050218E-02	9.55000000± 01 4.32000000€ 01 4.98953685E 05	7.99774067E-08 3.66374689E 02 7.20982867E 02 9.00983836E 09
3.50000000E 0 4.86545785E 0 1.59823821E 0	1.43488736E-02	9.48000000E 01 4.32000000E 01 4.96777804E 09	2.88638867E-08 3.67276211E 02 7.23238354E 02 2.80395099E 10
4.00000000E 0 4.82439914E 0 1.99148030E 0	5.134099846-03	9.40000000E 01 4.3200000E 01 4.94061055E 05	1.039162226-08 3.661793746 02 7.263094446 02 6.989374975 10
4.90000000F ( 4.788472766 ( 1.58594350F (	1.813280456-03	9.33000000E 01 4.32000000E 01 4.91852948E 05	3.68663129E-09 3.65084176E 02 7.24028986E 02 1.97749364E 11
9.00000000E ( 4.74741404E ( 1.57873125E (	05 6.37379063E-04	9.25000000 01 4.320000000 01 4.891005900 05	1.303169605-09 3.63990620E 02 7.32174756E 02 5.61843220E 11
6.000000000 4.0/042895E 1.50587841E	05 7.683624251-05	9.10000000 01 4.320000000 01 4.840/13286 05	1.987291756-10 3.618084276 02 7.381845005 02 4.650591176 12
7.00000000E 9.222039E 1.65576590E	05 1.024U4658E-05	8.99000000E 01 3.80000000E 01 9.449198/1E 09	1.88432819E-11 3.99632/99E 02 7.44344700E 02 3.99018617E 13
8.000000006 5.73855937k 1.73572584k	05 1.672897036-16	8.800000006 01 3.400000006 01 6.020078796 05	2.77686236E-12 3.57463724E 02 7.50661741E 02 2.70132753E 14
	06 1.14000000t-re 05 3.20894919t-n7 04 9.9483926t n9	8.65000000E 01 3.1000000E 01 6.52962311m U5	4.91382408F-13 3.75301213E 02 7.5/142394E 02 1.94083991E 15

INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 08 SURFACE GRAVITY= 3.75000000E 02 MOLECULAH MASS= 4.32000000E 01

INITIAL CONDITIONS

MEIGHT= 0 Pr SSURE= 4.00000000E 03 TEMPERATURE= 1.50000000E 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	NUMBER DENSITY	TEMPERATURE MOLECULAR MASS DENSITY SCALE MEIGHT	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
1.00000000E 07 6.73069667E 05 1.87979141E 04	2.48000000E-n5 7.02260585E-n8 2.11383874E ny	8.50000000E 01 2.80000000E 01 7.14723233E 05	9.82562974E-14 3.53145265E 02 7.63793846E 02 7.77348492E 15
1.10000000E 07 7.53838027E 05 1.98938424E 04	6.59000000E-n6 1.87751493E-n8 5.61701504E-08	8.50000000E 01 2.50000000E 01 8.05391969E 05	2.33118159E-14 3.50995877E 02 7.63793846E 02 3.27642363E 16
1.20000000 07 8.56634121E 05 2.12069072E 04	2.06000000E-n6 5.90506519E-n9 1.75584992E 08	8.50000000E 01 2.2000000E 01 9.20839866E 05	6.41269499E-15 3.48853050E 02 7.63793846E 02 1.19106530E 17
1.30000000E 07 9.42297533E 05 2.22419919E 04	7.13000000£-n6 2.05643347£-n8 6.07728638£ n9	8.50000000E 01 2.00000000E 01 1.01916490£ 06	2.01776325E-14 3.46716784E 02 7.63793846E 02 3.78534918E 16
1.40000000 07 9.91892140E 05 2.20198022E 04	2.680000000£-n7 7.77742453E-10 2.28430961E n7	8.50000000E 01 1.90000000E 01 1.07943558E 06	7.20508448E-16 3.44587079E 02 7.63793846E 02 1.06007618E 18
1.50000000E 07 1.10858533E 06 2.41248275E 04	1.09000000E-07 3.18281690E-10 9.29066220E 06	8.50000000E 01 1.7000000E 01 1.21390739E 06	2.62196024E-16 3.42463935E 02 7.63793846E 02 2.91306418E 18
1.60000000 07 1.32614520E 06 2.63860992E 04	4.68000000E-n8 1.37506579E-10 3.54301518t 06	9.5700000E 01 1.6000000E 01 1.46116739E 06	9.41073417E-17 3.40347392E 02 7.19829>13E 82 7.64902610E 18
1.7000000E 07 1.50158174E 06 2.80154544E 04	2.31000000±-n8 6.82952411±-11 1.56411003± n6	1.07000000 02 1.50000000 01 1.75346225E 06	3.69403505E-17 3.30237330E 02 6.80759529E 02 1.74785202E 19
1.8000000E 07 1.80142760E 06 3.14264820E 04	1.28000000£-n8 3.80000652£-11 7.89607333£ n5	1.10000000E 02 1.40000000E 01 2.09697141E 06	1.61422511E-17 3.36133669E 02 6.46066760E 02 3.56111665E 19

#### INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 08 SUPFACE GRAVITY= 3.75000000E 02 MOLECULAH MASS= 4.32000000E 01

INITIAL CONDITIONS

HEIGHT= 0 PHESSURE= 4.000000000 03 TEMPERATURE= 1.50300000E 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE TEMPERATURE COLUMNAR MASS MOLECULAR MASS NUMBER DENSITY DENSITY SCALE HEIGHT	DENSITY VISCOSITY(MIX.) VISCOSITY(KIM.)
1.9000000E 07 2.23763686E 06 3.42747626E 04	7.85000000E-09 1.31200000E 02 2.35003928E-11 1.30000000E 01 4.33485844E n5 2.51203878E 06	9.35510749E-18 3.34036969E 02 6.14778541E 02 6.57158180E 19
2.00000000E 07 2.66430008E 06 3.73999672E 04	5.20000000E-09 1.4420000E 02 1.56651688E-11 1.2000000E 01 2.61262562E 05 3.00985894E 06	5.20461893E-18 3.31946630E 02 5.86412187E 02 1.12671493E 20
2.50000000E 07 5.06915230E 06 5.15878373E 04	1.29000000t-09 2.17200000E 02 4.01127696E-12 9.50000000E 00 4.30297663E 04 5.91098077E 06	6.78614450E-19 3.21593351E 02 4.77810520E 02 7.04097179E 20
3.00000000E 07 8.62706791E 06 6.72994105E 04	5.96000000E-10 3.03500000E 02 1.91391187E-12 7.80000000E 00 1.42274367E 04 1.03889143E 07	1.84226362E-19 3.11404098E 02 4.04209271E 02 2.19409029E 21
3.50000000E 07 1.33328069E 07 8.36643509E 04	3.71000000E-10	7.42029298E-20 3.01376869F 02 3.50822132E 02 4.72787440E 21
4.00000000E 07 2.00245519E 07 1.02532384E 05	2.72000000E-10 5.1480000E 02 9.33048085E-13 5.7000000E 00 3.82797827E n3 2.57590117E 07	3.62222004E-20 2.91517666E 02 3.10360433E 02 8.56823795E 21
4.50000000E 07 2.72498802E 07 1.19608474E 05	2.200000000E-10 6.39100000E 02 7.80638772E-13 5.20000000E 00 2.49397997E U3 3.62596245E 07	2.15291466E-20 2.61820488F 02 2.78548747E 02 1.29382159E 22
5.00000000E 07 3.65833160F 07 1.38586583E 05	1.87000000E-10 7.75500000E 02 6.86774504E-13 4.7000000E 00 1.74702412E 03 5.03833329E 07	1.36309860E-20 2.72287336E 02 2.52868491E 02 1.85510052E 22
6.00000000E 07 5.86144216E 07 1.75421126E 05	1.510000000E-10	6.86975416E-21 2.53713107F 02 2.13890261E 02 3.11350677F 22

### INPUT CONSTANTS

PLANET RADIUS= 3.38188808E 08 SUPFACE GRAVITY= 3.75000000E 02 MDLECULAR MASS= 4.32088808E 01

### INITIAL CONDITIONS

MEIGHT= 0 PHESSURE= 4.000000000 03 TEMPERATURE= 1.500000000 02

HEIGHT PRESSURE SCALE HEIGHT SPOOD OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE Molecular mass Density scale Height	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
7.0000000E 07 8.86006034E 07 2.15674099E 05	1.31000000E-10 5.55567384E-13 6.59736235E 02	1.43860000E D3 3.60000000E DD 1.40907267E UR	3.94278730E-21 2.35794980F 02 1.85658822F 02 4.70882165E 22
6.00000000 07 1.19863722E 08 2.50876372E 05	1.19000000E-10 5.44540300E-13 4.68971065E 02	1.83640000E 03 3.40000000E 00 2.05719069E 08	2.64700935E-21 2.18532953E 02 1.64234461E 02 6.20454/83E 22
9.00000000E 07 1.50132769E 08 2.80131400E 05	1.11000000£-10 5.49703526£-13 3.52362227£ n2	2.28230000E 03 3.20000000E 00 2.93669395E 08	1.87184479E-21 2.01927029F 02 1.47400600E 02 7.87461659E 22
1.00000000E 08 2.04691805E 08 3.27807791E 05	1.05000000t-10 5.64585320t-13 2.74679993t 02	2.76950000E 03 3.00000000E 00 4.12715510E 08	1.36797699E-21 1.65977205E 02 1.33808679E 02 9.76151535E 22
2.00000000E 08 7.83144272E 08 6.41210373E 05	7.900000000±-10 1.26269601±-11 5.78716022£ 62	9.89010000E 03 2.80000000E 00 4.69401810E 09	2.69001095E-21 6.25645440E 01 7.08084807E 01 2.63227481E 22
3.0000000E 08 1.77578315E 09 9.65549665E 05	7.21000000E-11 1.51406454E-11 2.50847748E 01	2.08240000E 04 2.60000000E 00 1.39839648E 11	1.06271478E-22 4.76201630E 00 4.6/981705E 01 4.50701989E 23
4.00000000 98 3.13224134£ 09 1.28235202£ 06	6.900000000t-11 5.40942513E-12 1.41544203t 01	3.53180000E 04 2.5000000E 00 9.34467628E 10	5.87438769E-23 1.2569622ZE 01 3.74703342E 01 6.3785940ZE 23
5.00000000 08 5.12899054E 09 1.640950956 06	6.720000000-11 7.815191966-13 9.15056185t 00	5.32060000E 04 2.30000000E 00 2.23680715E 10	3.493864895-25 8.598736176 01 3.052848926 01 8.737741776 23
6.00000000E 08 7.85131807E 09 2.03025063E 06	6.62000000t-11 2.94202302t-13 6.44962323t NU	7.43640000E 04 2.10000000E 00 1.50846442E 10	7.24845474E-23 2.25015235E 02 2.58228544E 01 1.14847117E 24

#### INPUT CONSTANTS

PLANET RADIUS= 3.38100000E 08 SURFACE GRAVITY= 3.75000000E 02 MOLECULAR MASS= 4.32000000E 01

INITIAL CONDITIONS

HEIGHT= 0 PRESSURE= 4.000000000 03 TEMPERATURE= 1.50000000E 02

HEIGHT PRESSURE SCALE HEIGHT SPEED OF SOUND	PRESSURE COLUMNAR MASS NUMBER DENSITY	TEMPERATURE HOLECULAR MASS DENSITY SCALE HEIGHT	DENSITY GRAVITY VISCOSITY(MIX.) VISCOSITY(KIM.)
7.0000000E 08 1.09412938E 10 2.39670174E 06	6.54000000E-11 1.52215772E-13 4.80084069E 00	9.86960000E 04 2.00000000E 00 9.54952688E 09	1.59396140E-23 4.29653241E 02 2.24148503E 01 1.40623545E 24
8.00000000E 08 1.55553453E 10 2.85588100E 06	6.49000000E-11 9.27273495E-14 3.72811673E 00	1.26123000E 05 1.80000000E 00 8.32367918E 09	1.11401879E-23 6.99901381E 02 1.98284365E 01 1.77990144E 24
2.0000000E 08 2.15973707E 10 3.37507328E 06	6.46000000E-11 6.23696817E-14 2.98912205E 00	1.56577000€ 05 1.60000000€ 00 7.85560045€ 09	7.93951806E-24 1.03575966E 03 1.77959687E 01 2.24144194E 24
1.00000000E 09 2.80841618E 10 3.83981574E 06	6.43000000E-11 4.47388982E-14 2.45186452E 00	1.90000000E 05 1.50000000E 00 7.32768928E 09	6.10545787E-24 1.43722806E 03 1.61550690E 01 2.64600450E 24

END OF PROBLEM

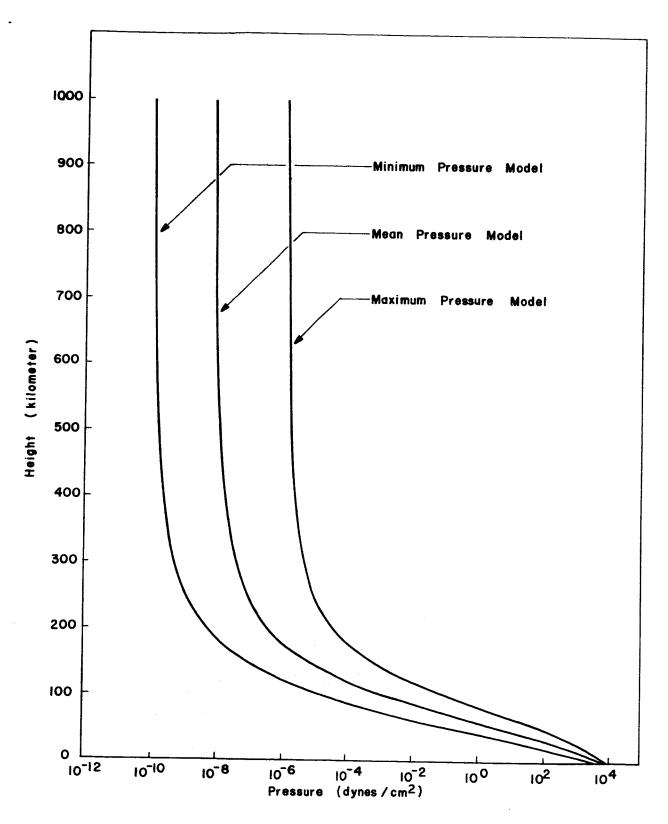


Figure 1: Variation of Pressure with Altitude for Three Atmospheric Models

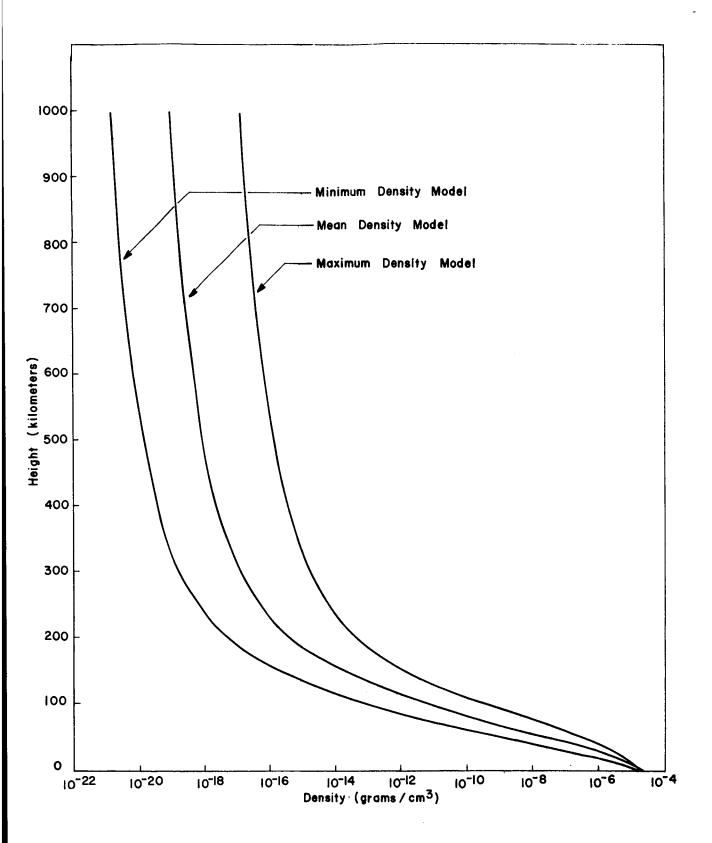


Figure 2: Variation of Density with Altitude for Three Atmospheric Models

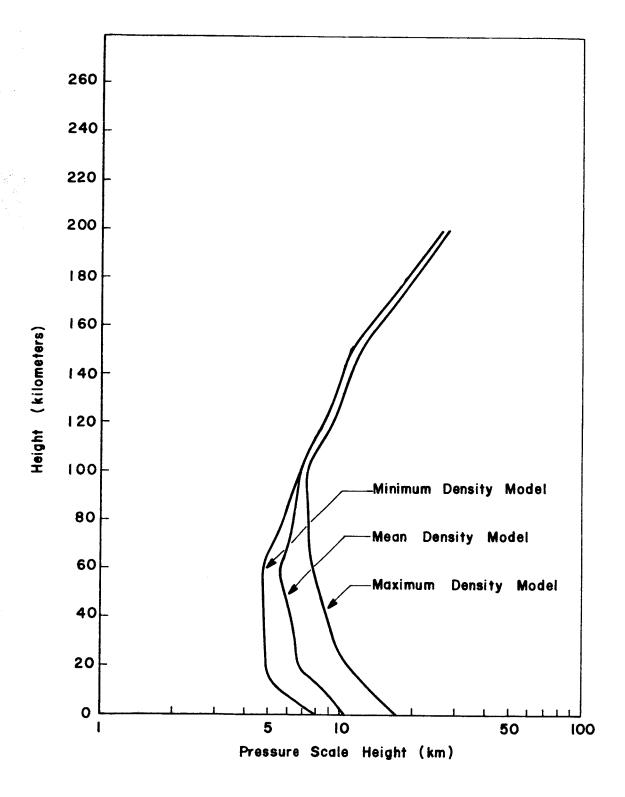


Figure 3: Variation in the Pressure Scale Height for Three Atmospheric Models

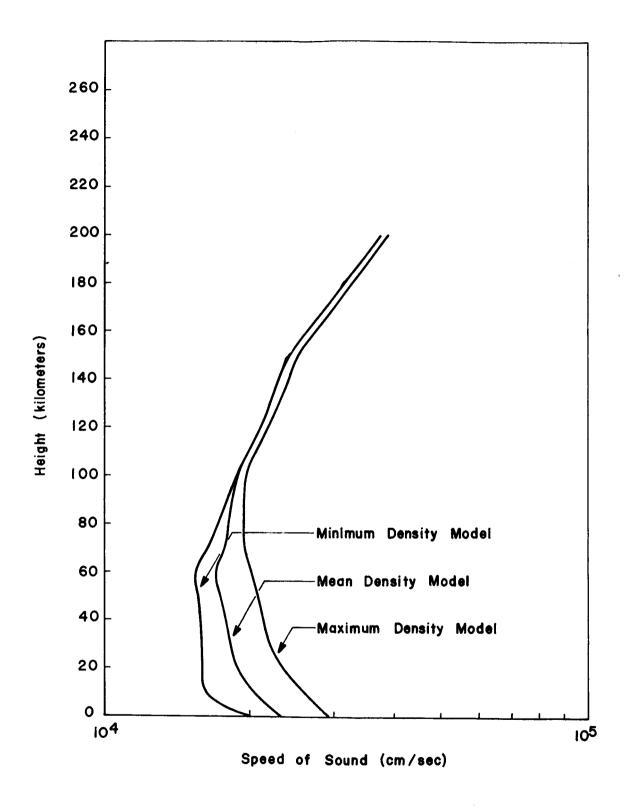


Figure 4: Variation in the Speed of Sound for Three Atmospheric Models

#### **REFERENCES**

- 1. Owen, Robert B., 'The Martian Environment,' NASA TM X-53167, Marshall Space Flight Center, Huntsville, Alabama, November 19, 1964.
- 2. Johnson, Francis S., 'Atmosphere of Mars,' SCIENCE, Vol 150, pp 1445-1448, December 10, 1965.
- Kliore, Arvydas, et al., "Occultation Experiment: Results of the First Direct Measurement of Mars' Atmosphere and Ionosphere," SCIENCE, Vol. 149, pp 1243-1248, September 10, 1965.
- Smith, Edward J., et al., 'Magnetic Field Measurements Near Mars,'' SCIENCE, Vol. 149, pp 1241-1242, September 10, 1965.
- 5. O'Gallagher, J. J., and J. A. Simpson, 'Search for Trapped Electrons and a Magnetic Moment at Mars by Mariner IV,' SCIENCE, Vol. 149, pp 1233-1239, September 10, 1965.
- Van Allen, J. A., et al., 'Absence of Martian Radiation Belts and Implications Thereof," SCIENCE, Vol. 149, pp 1228-1233, September 10, 1965.
- 7. Kaplan, L. D., G. Münch, and H. Spinrad, 'An Analysis of the Spectrum of Mars," THE ASTROPHYSICAL JOURNAL, Vol. 139, No. 1, January 1, 1964.
- Spinrad, H., G. Munch, and L. D. Kaplan, 'The Detection of Water Vapor on Mars,' ASTROPHYSICAL JOURNAL, pp 1319-1321, May 15, 1963.
- 9. Kern, L. C., and G. F. Schilling, 'Modat: A Computer Program for the Construction of Model Atmosphers," Memorandum RM-4204-PR, The Rand Corporation, July 1964.
- Brakaw, Richard, S., 'Alignment Charts for Transport Properties, Viscosity, Thermal Conductivity, and Diffusion Coefficients for Nonpolar Gases and Gas Mixtures at Low Density," NASA TR R-81, Lewis Research Center, Cleveland, Ohio, 1961.
- Arnold, James O., Victor H. Reis, and Henry T. Woodward, "Studies of Shock-Layer Radiation of Bodies Entering Planetary Atmospheres," AIAA JOURNAL, Vol. 5, No. 11, pp 2019-2025, November 1965.

# THE ATMOSPHERE OF MARS: A DERIVATION OF ENGINEERING AND DESIGN PARAMETERS

By W. T. Roberts and George S. West

The information in this report has been reviewed for security classification. Review of any information concerning Department of Defense or Atomic Energy Commission programs has been made by the MSFC Security Classification Officer. This report, in its entirety, has been determined to be unclassified.

This document has also been reviewed and approved for technical accuracy.

Robert E. Smith

Chief, Space Environment Branch

William W. Vaughan

Chief, Aerospace Environment Division

E. D. Geissler

Director, Aero-Astrodynamics Laboratory

## DISTRIBUTION

DEP-T	R-AERO (continued)
Dr. Rees	Mr. Dahm (2)
	Mr. Dalton
R-DIR	Mr. McNair
Mr. Weidnes	Mr. Lewis
Dr. McCall	Mr. Daniels
Dr. Johnson	Mr. Belew
	Mr. Roberts (50)
R-RP	Mr. O. Jean
Dr. Stuhlinger (3)	Mr. Lindberg (2)
Dr. Shelton	Mr. Thomae (4)
Mr. Urban	Mr. Hasseltine
Mr. Burrell	Mr. Wilson
Mr. Downey	
Dr. Dozier	R-AS
	Mr. Williams (5)
R-P&VE	Mr. Huber
Dr. Lucas	Mr. Waggoner
Mr. Gause	Mr. Woodcock
Dr. Hellebrand	Mr. Paul
Mr. Goerner	Mr. Carter
Mr. Riehl	Mr. Gradecak
Mr. Kingsbury	Mr. Spears
Mr. Cataldo	
Mr. Shannon	MS-IP
	MS-H
R-ASTR	MS-T (6)
Dr. Haeussermann	MS-IL (8)
Mr. Currie	CC-P
Mr. Boehm	HME-P
Mr. Hoberg	
1-DIR	
Gen. 0'Connor (4)	
R-AERO	
Dr. Geissler	
Mr. W. Vaughan (2)	
Mr. R. Smith	
Mr. O. Vaughan	
Mr. O. Smith	
Mr. J. Scoggins	
Mr. Horn (2)	
Mr. Baker (2)	

### EXTERNAL DISTRIBUTION

```
NASA Headquarters
Federal Office Bldg. 6
Washington 25, D. C.
Attn: Technical Information Division (2)
Scientific and Technical Information Facility (25)
P. 0. Box 5700
Bethesda, Maryland
Attn: NASA Rep. (S-AK/RKT)
NASA
Office of Manned Space Flight
Federal Office Bldg. 6
Washington 25, D. C.
Attn: Director (4)
       Mr. E. E. Christensen
       Mr. E. Z. Gray
       Dr. E. J. McLaughlin
NASA
Office of Space Science and Applications
Federal Office Bldg. 6
Washington, D. C.
Attn: Dr. Newell (2)
       Dr. John E. Naugle
       Dr. M. Tepper
       Dr. Schmerling
       Dr. Henry J. Smith
NASA
Office of Advanced Research and Technology
Washington 25, D. C.
Attn: Dr. M. Adams
       Mr. Keller
       Mr. Reetz
       Mr. Rhode
       Mr. Charak
```

## EXTERNAL DISTRIBUTION (Continued)

```
Ames Research Center
Moffett Field, California
Attn: Dr. John R. Spreiter
       Dr. Barrette S. Baldwin
       Library (2)
NASA
Manned Spacecraft Center
Houston 1, Texas
Attn: Mr. R. Thompson
       Mr. J. Modisette
       Mr. D Robbins
       Mr. J. Harris (3)
NASA
Jet Propulsion Laboratory
4800 Oak Grove Drive
Pasadena, California
Attn: Mr. A. J. Beck
       Technical Library (2)
```

ORA (RRRS/Dr. Paul D. Jose)

Holloman Air Force Base, New Mexico

NASA

NASA

NASA

Goddard Space Flight Center

Dr. N. Ness

Library (2)

Langley Research Center Langley Field, Virginia

> Dr. Foelsche Mr. Crouch Library

Attn: Director

Dr. F. B. McDonald

Mrs. H. H. Malitson

Greenbelt, Maryland Attn: Dr. Clark

## EXTERNAL DISTRIBUTION (Continued)

National Research Council
Radio and Electrical Engineering Division
Ottawa 2, Canada
Attn: Dr. Peter D. Millman

Laboratory of Tree-Ring Research University of Arizona Tuscon, Arizona 85721 Attn: Dr. Harold C. Fritts

Smithsonian Astrophysical Observatory 60 Garden Street Cambridge 38, Massachusetts 02138 Attn: Dr. Fred Whipple Dr. Charles Lundquist Dr. Gerald Hawkins

Aeronomy Laboratory
ITSA
Department of Commerce
Boulder, Colorado
Attn: Dr. Franklin E. Rooch
Library (2)

Dr. Luigi Jacchia

Bellcomm Inc. 1100 17th Street, N. W. Washington, D. C. 20036 Attn: Dr. G. T. Orrok

Air Force Cambridge Research Laboratories Cambridge, Massachusetts Attn: Dr. E. J. Chernosky Library (2)

Dr. E. J. Öpik University of Maryland College Park, Maryland

Dr. S. Fred Singer University of Miami Miami, Florida

## EXTERNAL DISTRIBUTION (Continued)

Dr. Hector R. Rojas IIT Research Institute Astro-Sciences Center 10 West 35th Street Chicago, Illinois

Dr. Robert F. Rolsten Research Institute University of Dayton Dayton, Ohio 45409

Dr. A. E. S. Green 3535 North West 7th Place Gainesville, Florida 32601

Dr. James I. Vette
Aerospace Corporation
Box 95085
Los Angeles, California 90045

Dr. Carl McIllwain University of California at San Diego La Jolla, California

Dr. J. B. Blizard University of Denver Denver, Colorado

Hughes Aircraft Company Aerospace Group Space Systems Division El Sequndo, California Attn: Dr. Samuel Sabaroff

Dr. Seymour L. Hess Department of Meteorology The Florida State University Tallahassee, Florida 32301

Dr. Robert Jastrow Goddard Institute for Space Studies 2880 Broadway New York, New York 10025 Dr. P. H. Staher
Department of Physics
Potchefstroom University for C.H.E.
Potchefstroom, South Africa

Roy O. West Library
Depauw University
Greencastle, Indiana
Attn: Dan Smith
Circulation Librarian

Dr. Raymond Davis, Jr. Brookhaven National Laboratory Upton, L. I., New York 11973

Dr. J. A. Van Allen Department of Physics and Astronomy State University of Iowa Iowa City, Iowa

Dr. John C. Noyes Head, Geo-Astrophysics Boeing Scientific Research Lab. P. O. Box 3981 Seattle, Washington 98124